

Sintered Icy Regolith Simulants and their Implications for Lunar Subsurface Modification on Geologic Timescales

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Introduction

- Research questions:
 - How do we expect water ice to interact with lunar regolith?
 - How can we best produce icy lunar regolith simulants on Earth to replicate the expected properties?
- Motivation of research: Facilitate exploitation of lunar ice as a resource

Acknowledgements

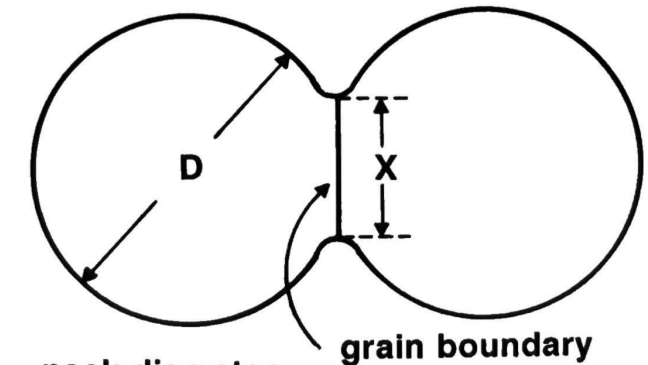
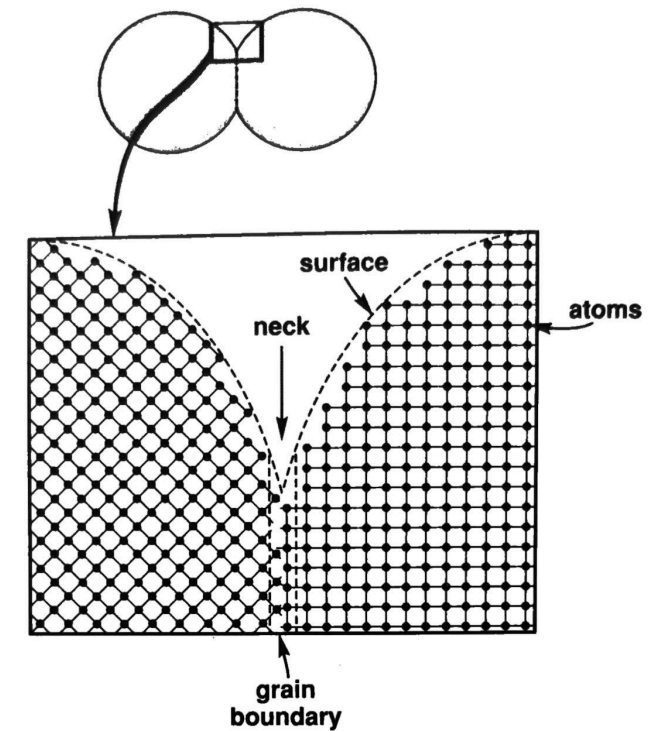
- CRREL (US Army's Cold Regions Research and Engineering Lab)
 - Lab facilities, testing support
- Colorado School of Mines Space Resources Department
 - Funding



Sintering

Sintering Basics

- Bonding of two or more solid-state particles by mass transport to a neck region
 - Degree of sintering usually defined by neck growth, neck size ratio
- Surface area reduces
- Relevant temperatures: $T > 50\%$ of melting point, or close to a phase change

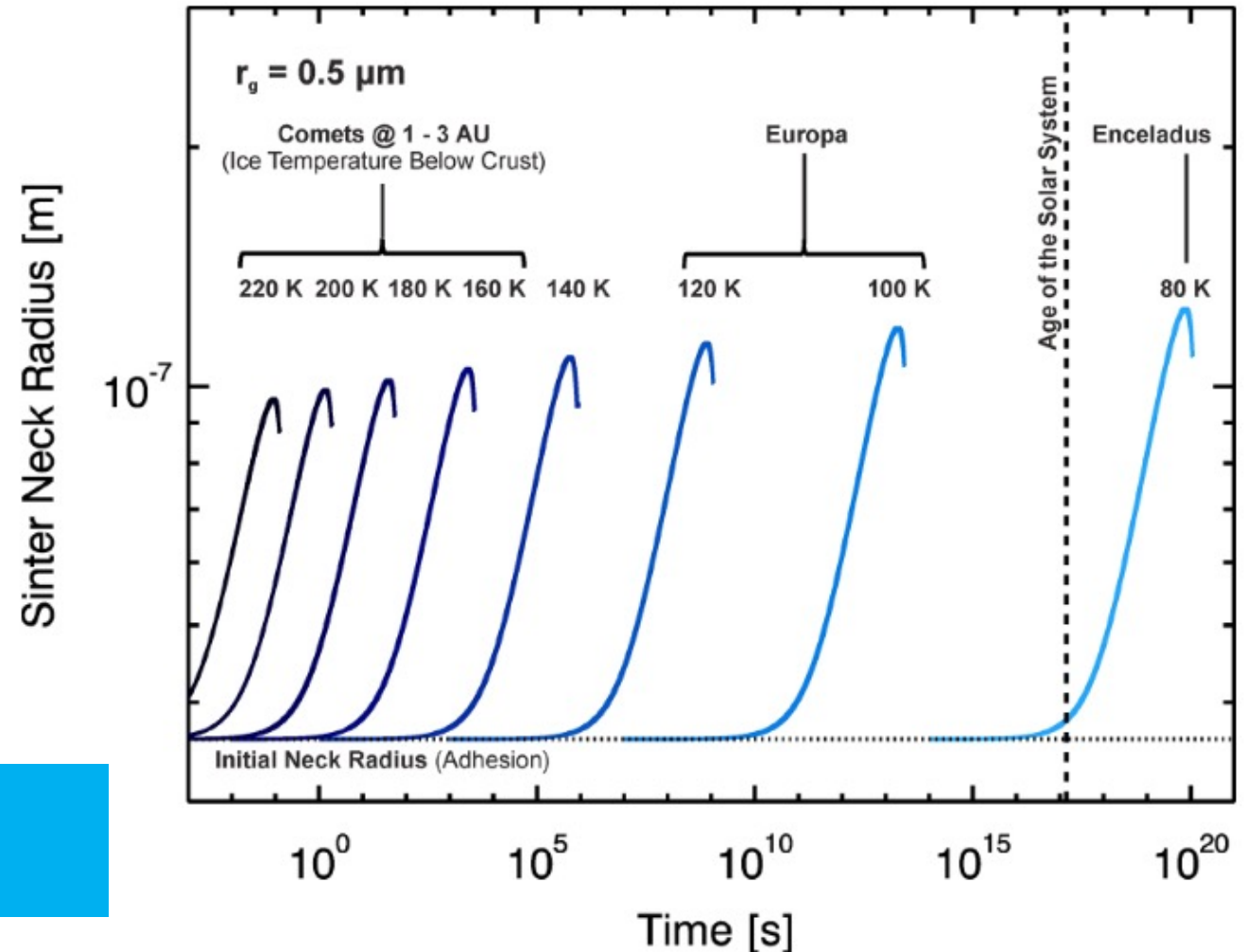


X = neck diameter
 D = particle diameter
 X/D = neck size ratio

Ice Sintering in Planetary Science

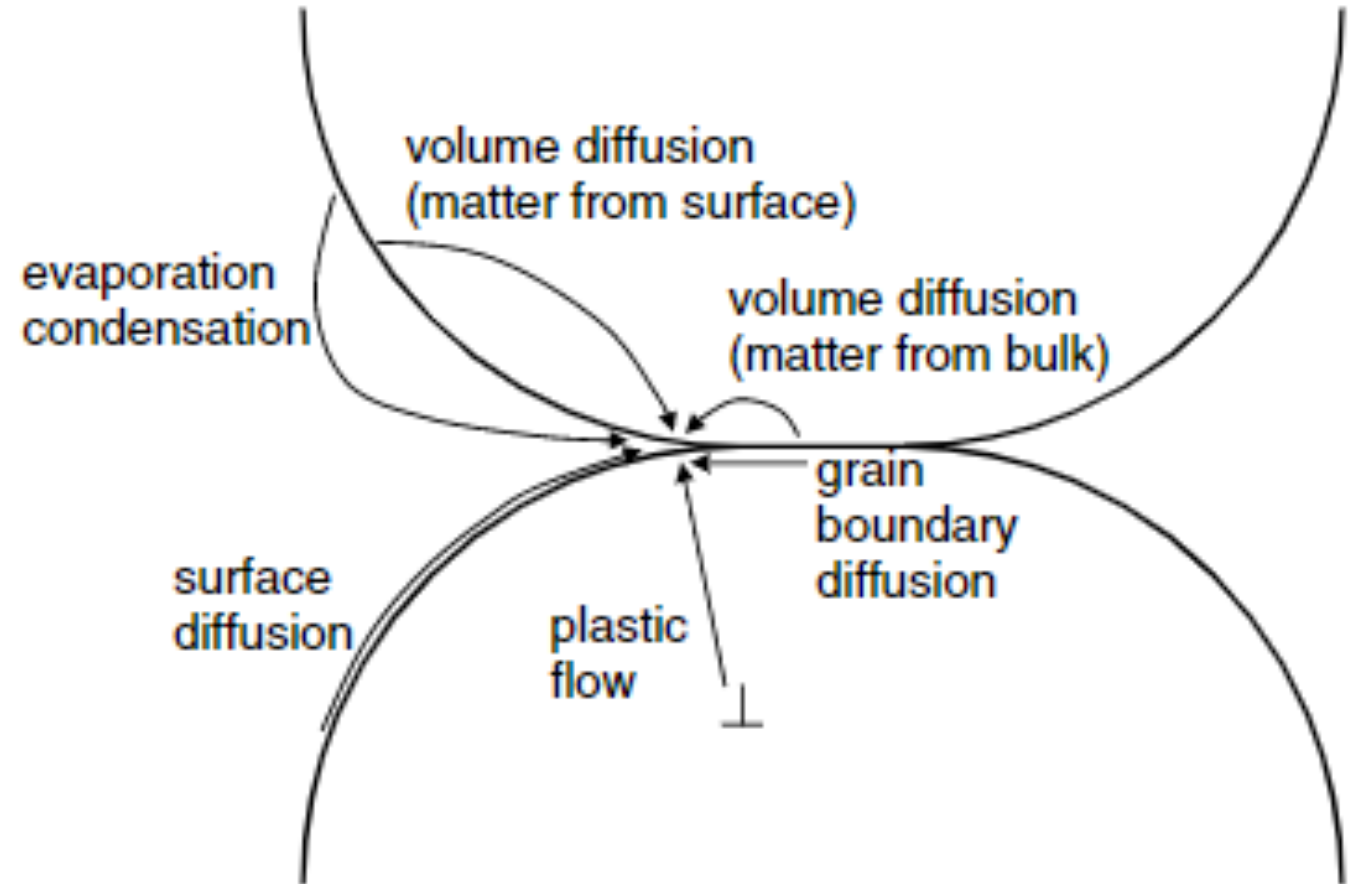
- Cryogenic Ice Vacuum Sintering Experiments
 - 120 K: 32 years
 - 100 K: 320,000 years
 - 80 K: >age of the solar system
- Ice sintering shown to occur at temperatures and pressures relevant to lunar PSRs

Ice is Hot!



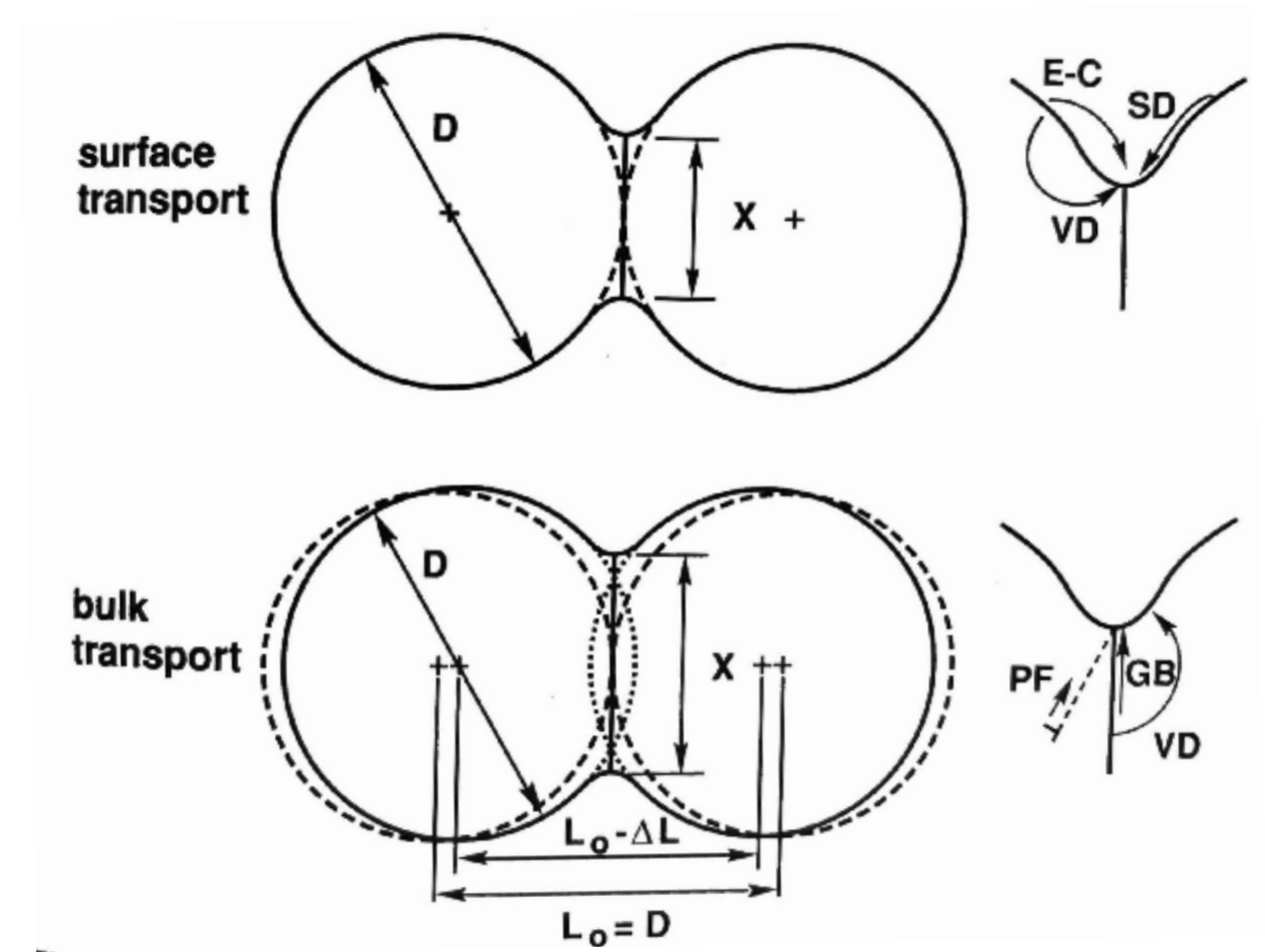
Sintering Mass Transport Mechanisms

- Two categories of mass transport: surface and bulk transport
- Surface Transport:
 - sublimation-deposition
 - surface diffusion
 - volume diffusion
- Bulk Transport:
 - plastic flow
 - volume diffusion
 - creep



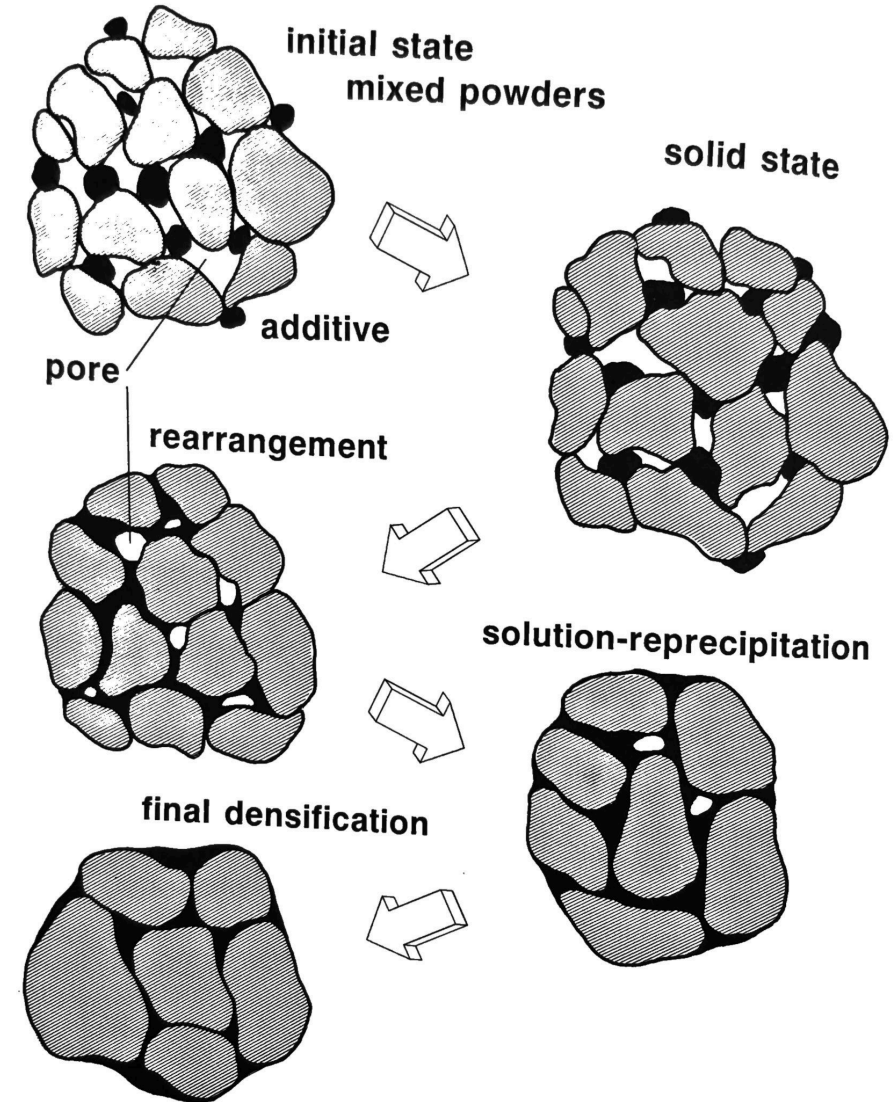
Surface vs. Bulk Transport

- Surface Transport: no densification
 - Reduction in total surface area, no change in bulk density
- Bulk Transport: Significant densification
 - Pressure-assisted sintering often used to drive bulk transport and densification in industrial processes



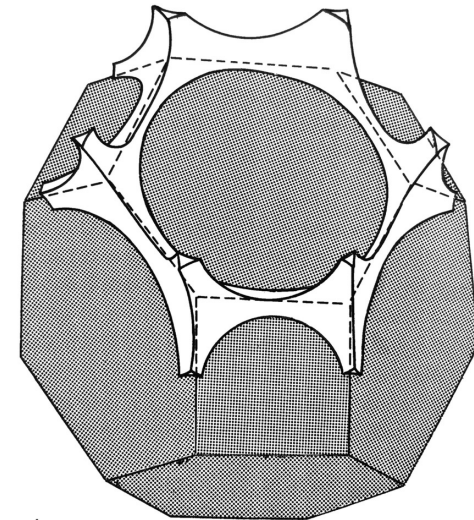
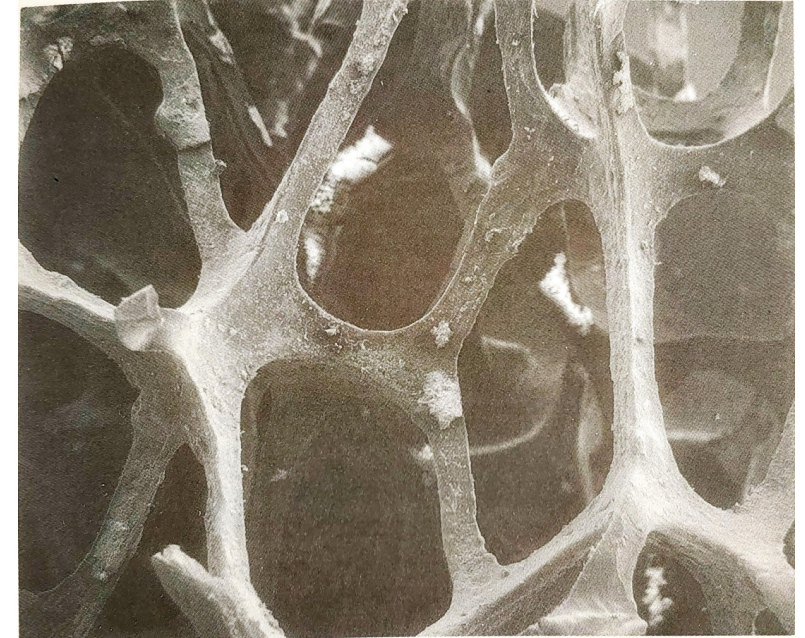
Liquid Sintering

- At least one phase is liquid
 - Characteristic of very high sintering rates
- Liquid phase lubricates particles, allows rearrangement and advanced densification



Liquid Sintering Contd.

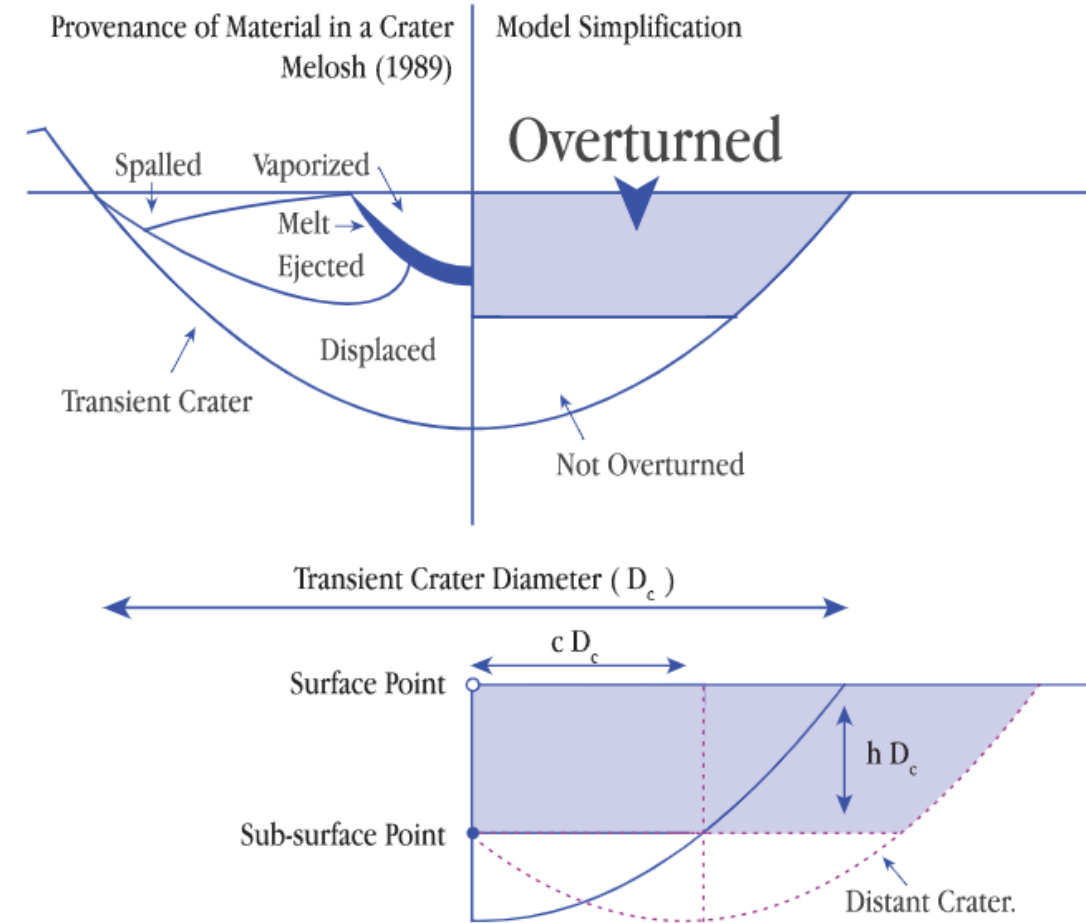
- Liquid penetrates pore space and solidifies in a continuous matrix
- Images of the solidified liquid phase shown on the right



Lunar Water Emplacement and Modification

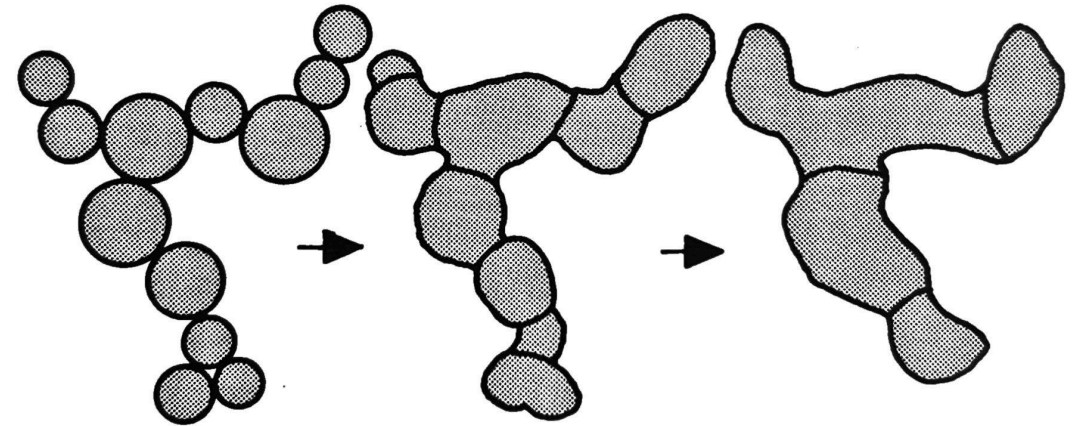
Water Emplacement and Modification Overview

- Collisional water atmosphere formed primarily by asteroid impacts
 - Water deposits in cold traps at the poles
- Impact gardening over billions of years causes mixing
 - Possible desiccation in top layers

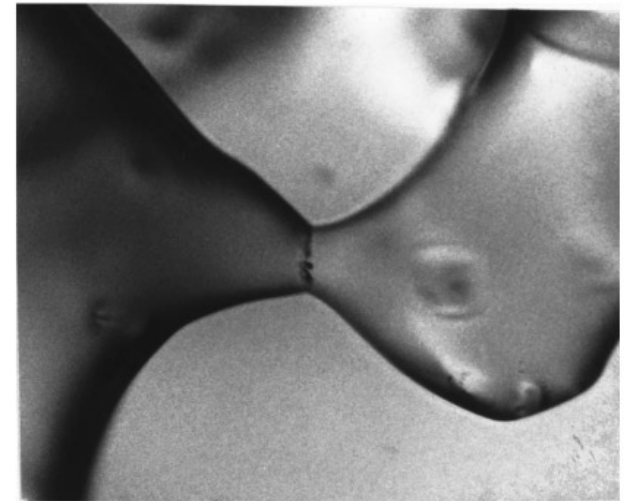
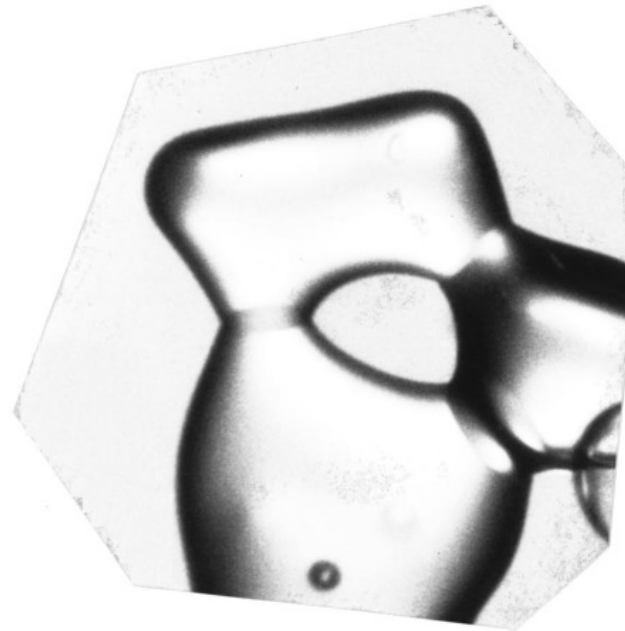


Ice Sintering

- Extensive literature exists on the sintering of dry snow
 - Water has a high vapor pressure and is generally found in nature near its triple point, so surface-transport is the primary driver
 - Maintains bulk density through sintering, but particles combine and coarsen
- Bulk transport occurs with application of pressure



microstructure coarsening



Icy Simulants

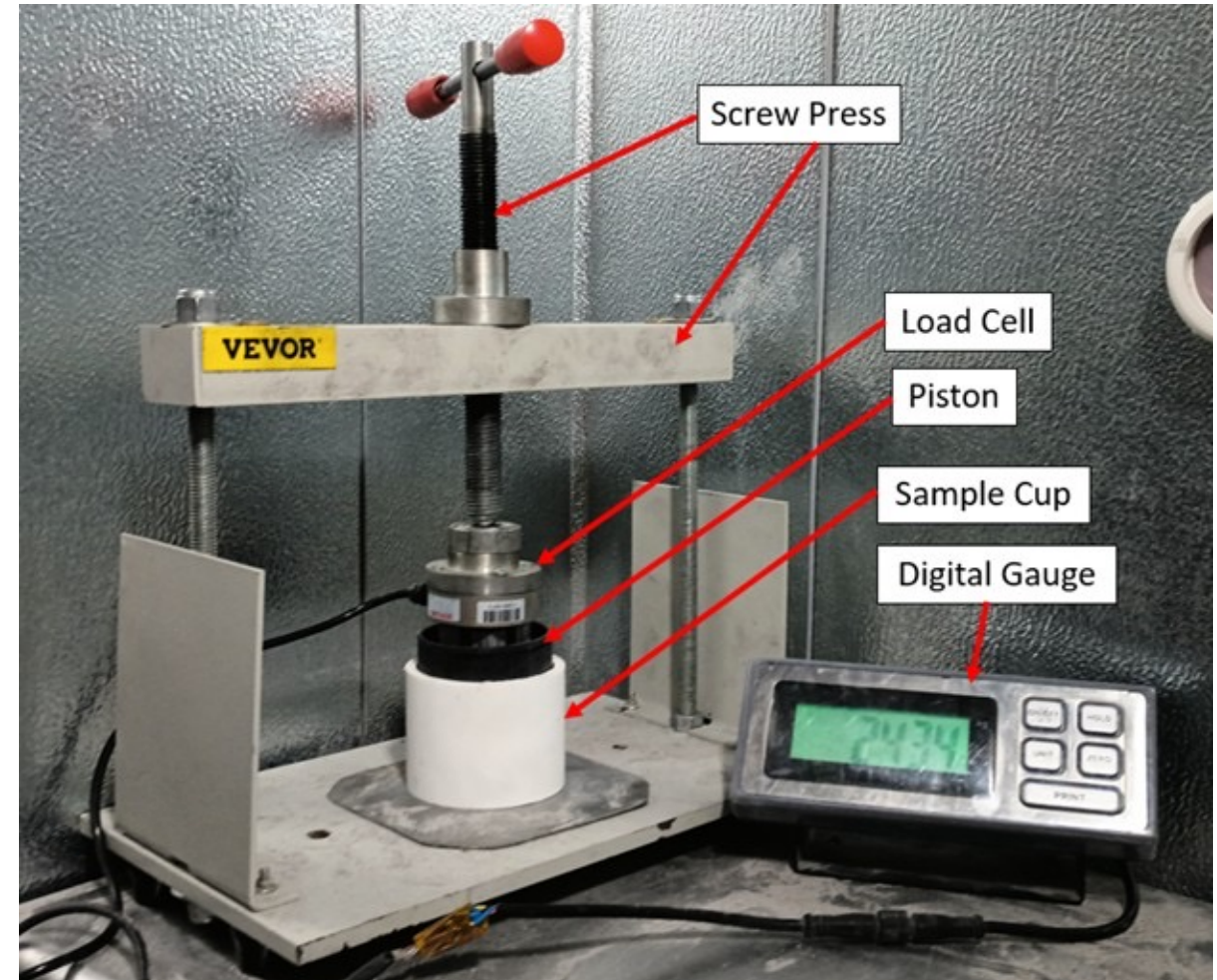
Wet Mix Icy Regolith Simulant

- Oldest, most common method in literature
- Mix liquid water with regolith simulant
 - Sometimes compact
- Freeze mixture
- Characteristics: Very high strength, limestone or concrete
- Requires aggressive excavation methods
 - Jackhammers, impact drills, etc.



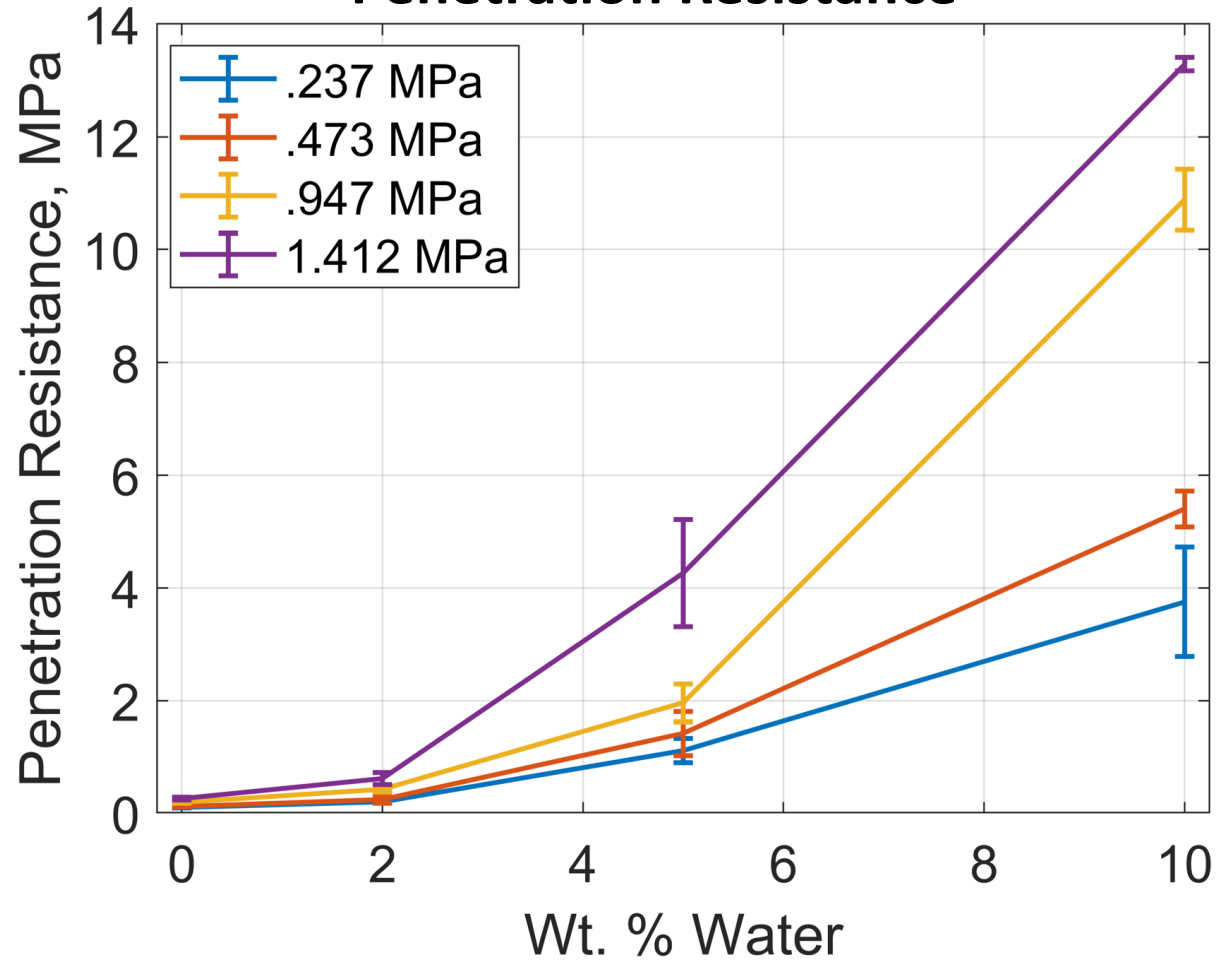
Pressure Sintered Icy Lunar Regolith Simulant (PSS)

- Mixture of sieved ice grains and regolith simulant placed in cup
- Cup pressed with piston for 10 minutes
- Ice grains sinter the mixture together
- At -10°C , melting pressure 150 MPa, Yield Strength ~ 1.5 MPa

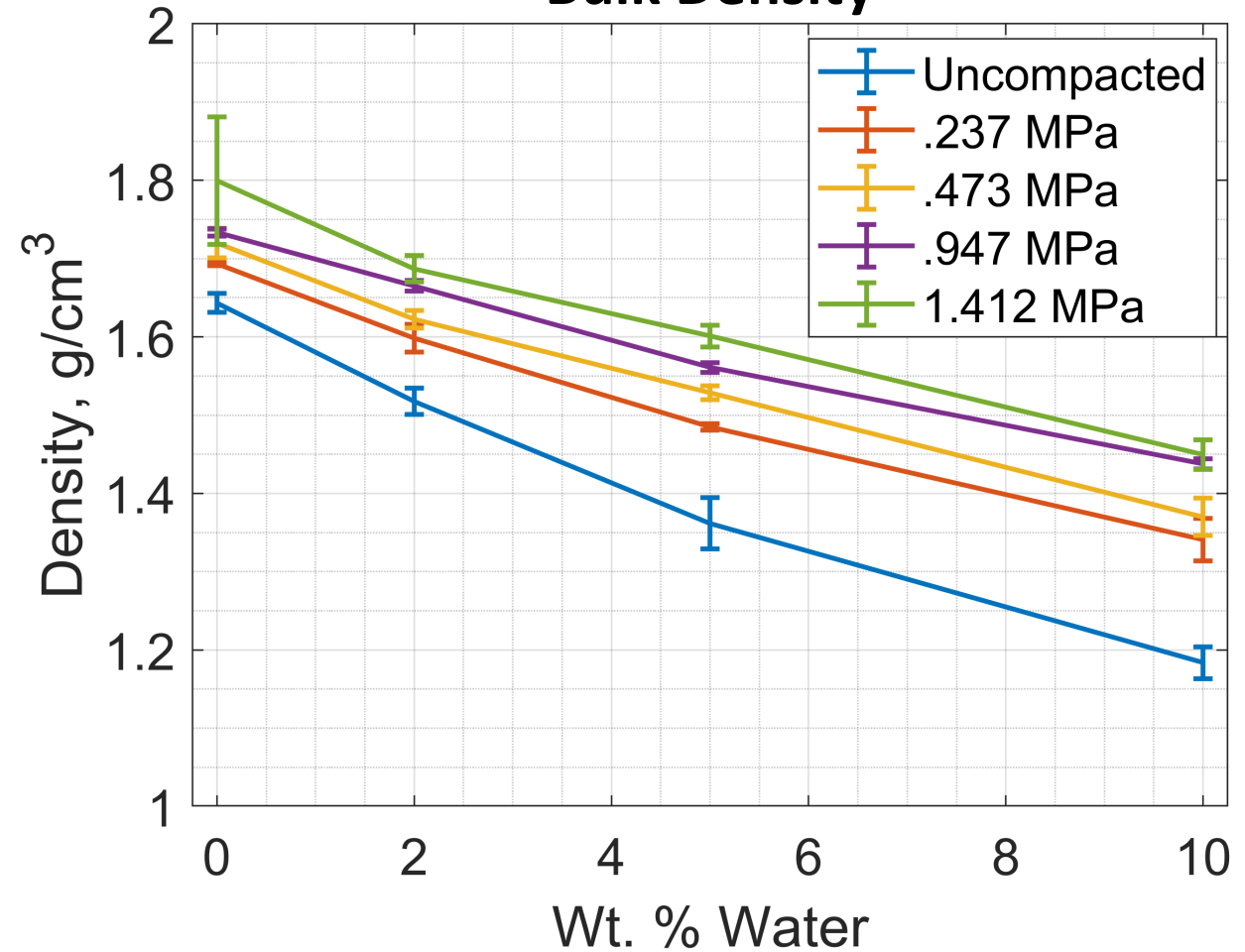


PSS Behavior

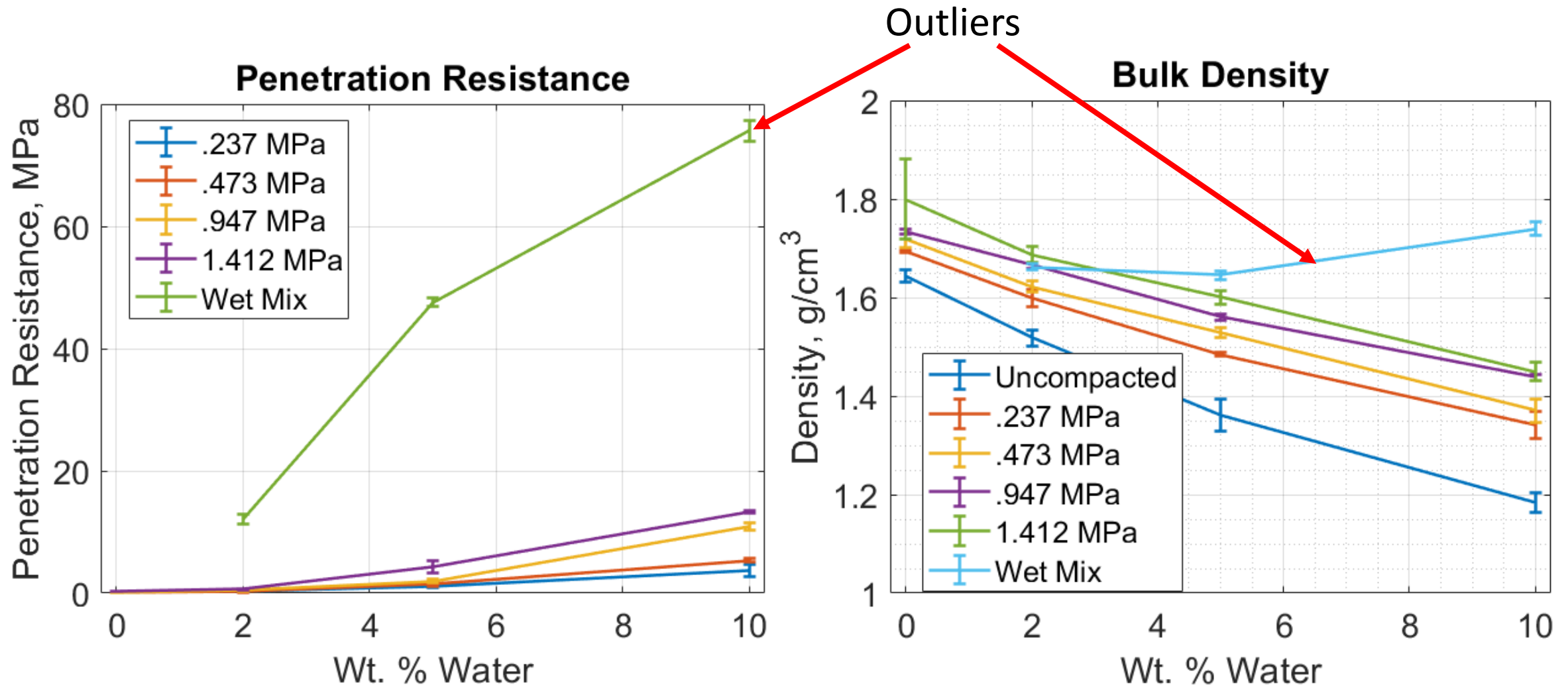
Penetration Resistance



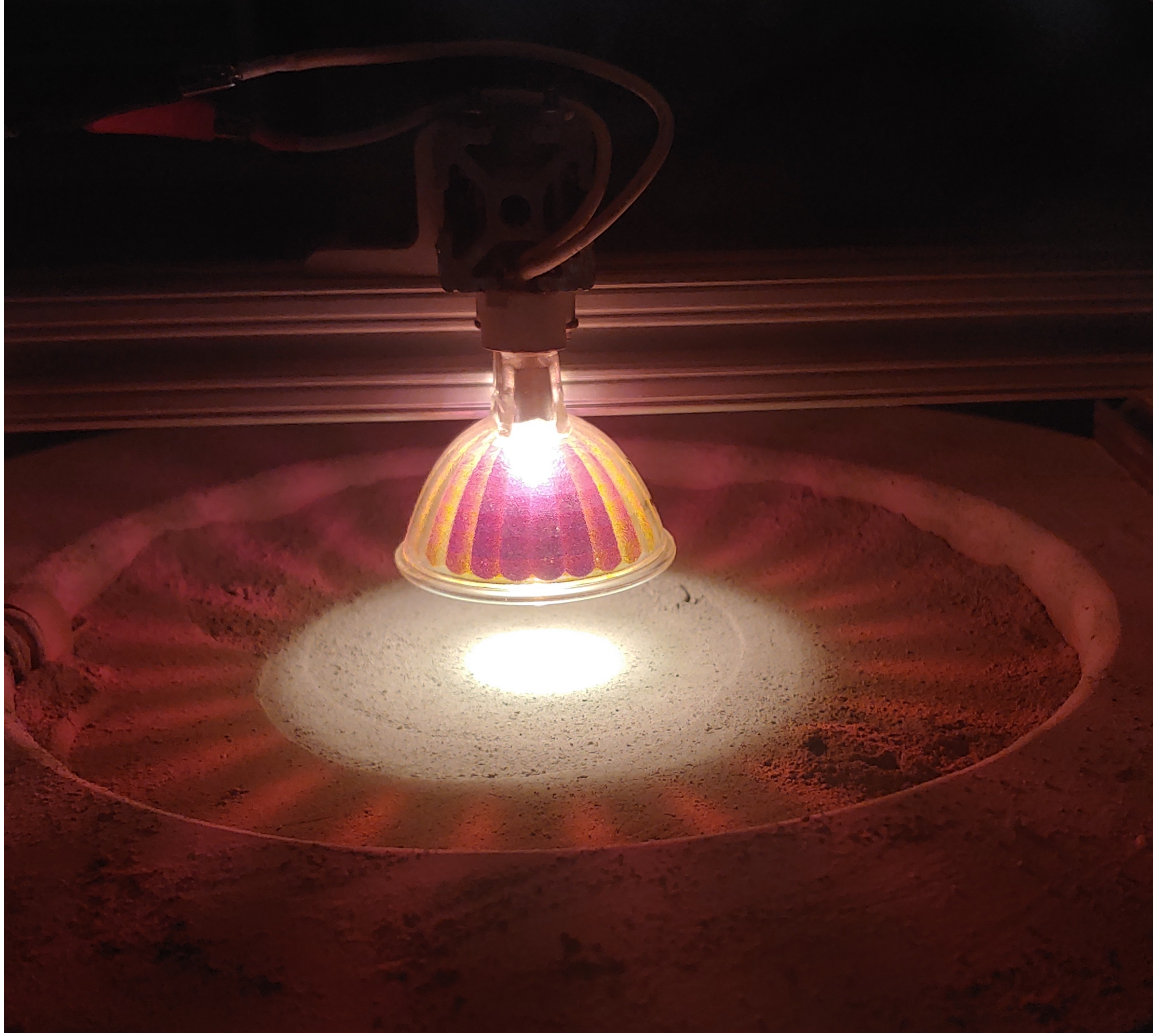
Bulk Density



PSS Behavior Compared to Wet Mix

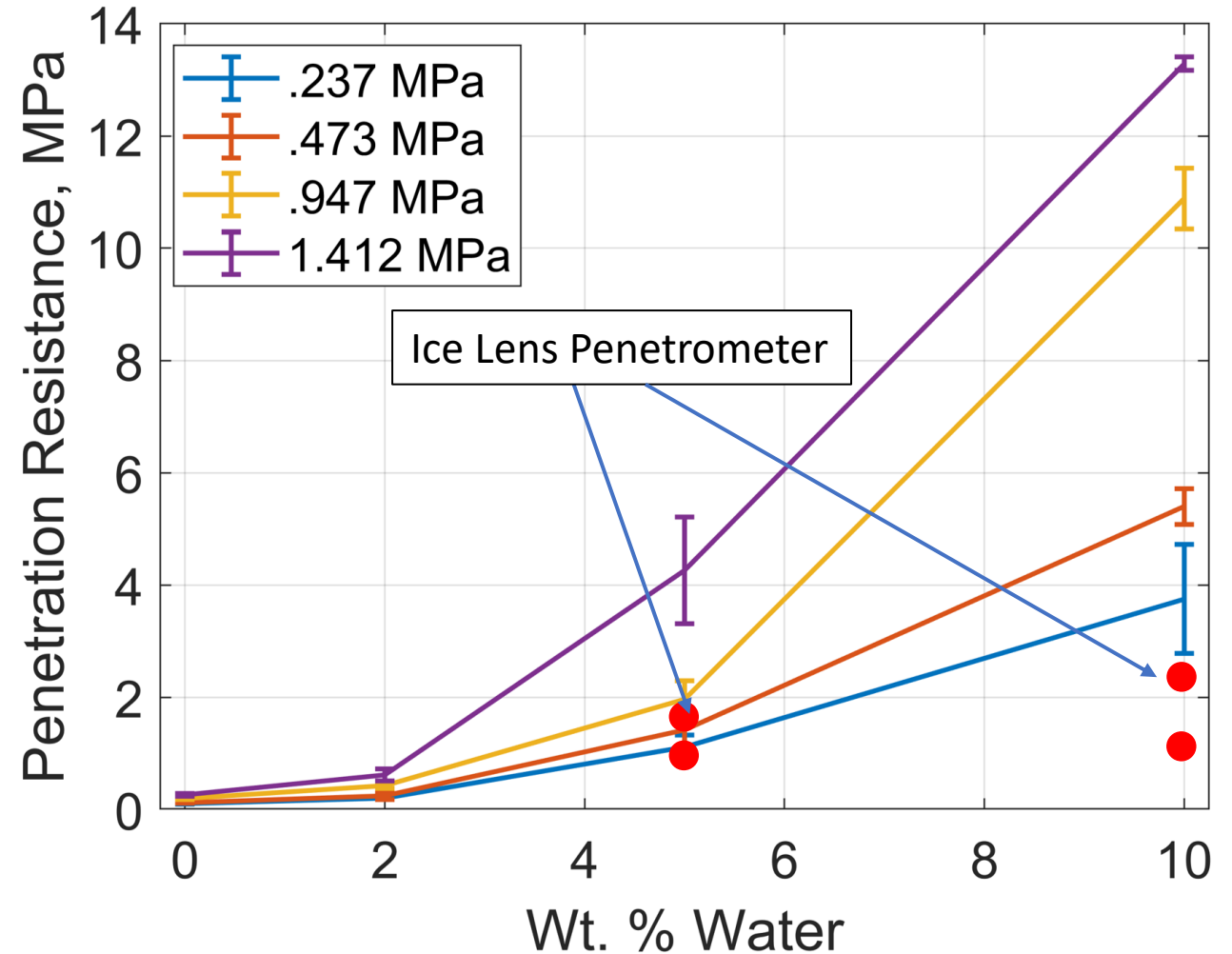


Vacuum Sintered Icy Regolith Simulant (VSS)



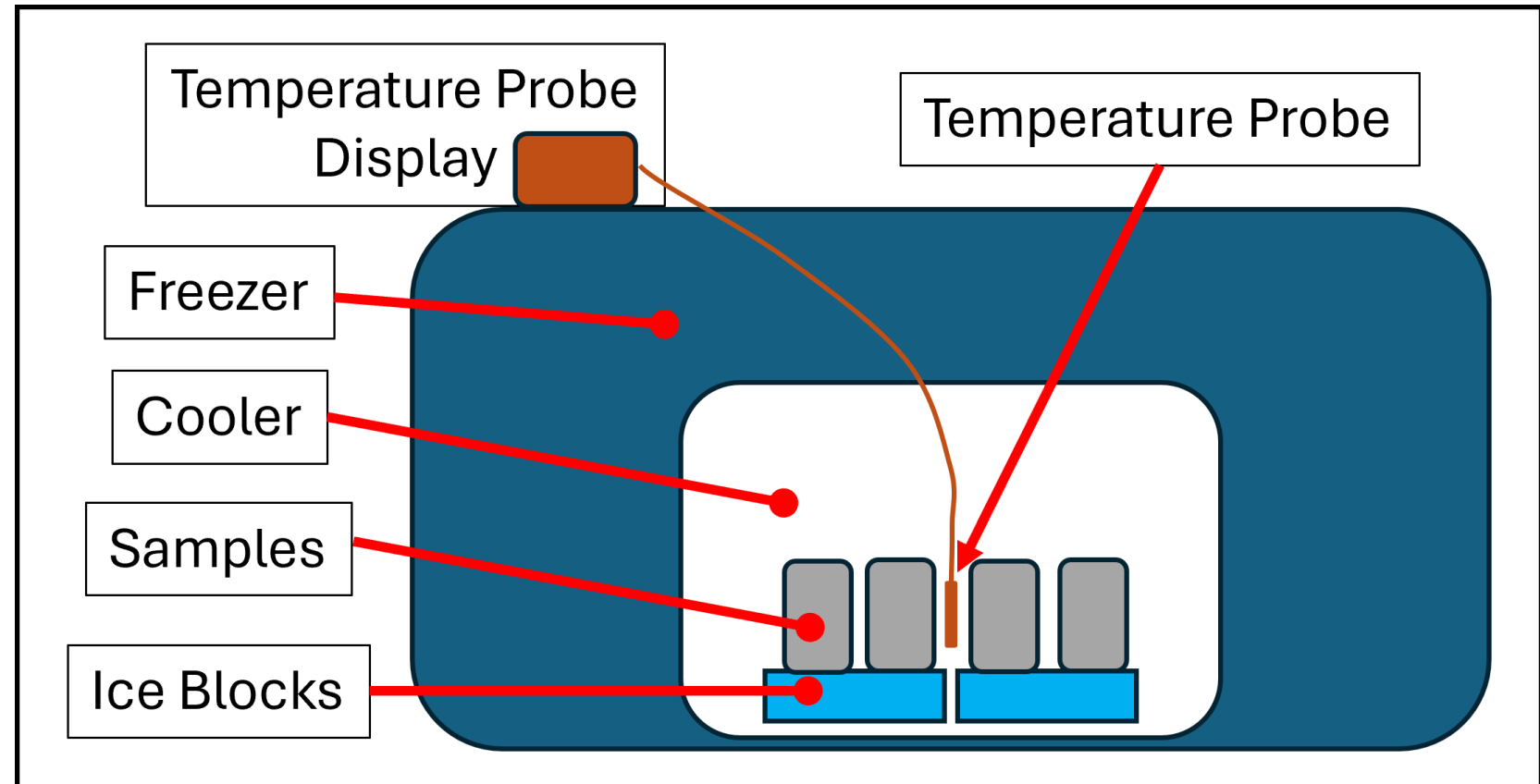
Results: Thermal Sintering

- Variable ice lens thicknesses observed
- Penetration resistance similar to lower-strength PSS
- 40-80x weaker than wet-mix
- No change in ice content observed in ice lens
 - Sintering much faster than diffusion, desiccation

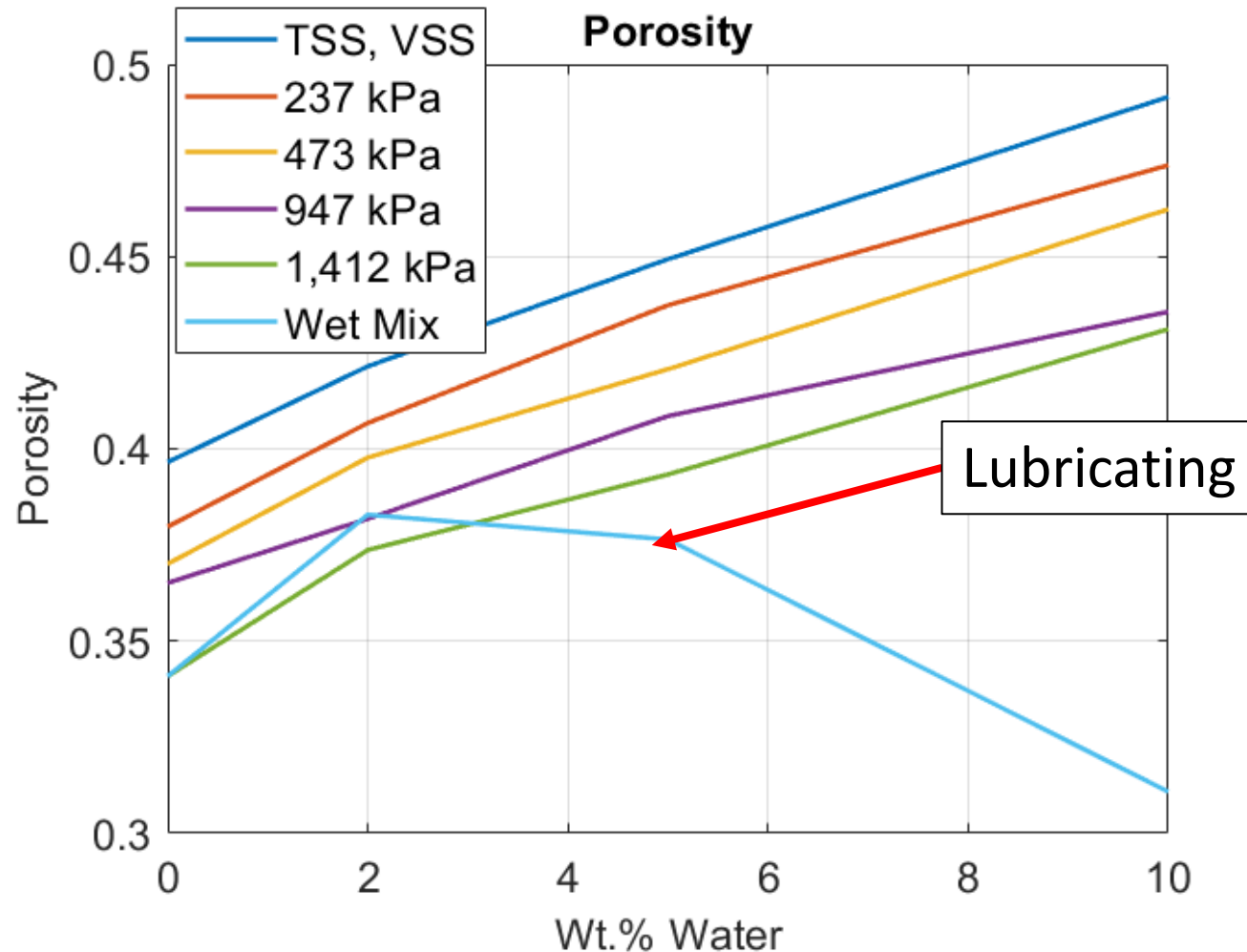


Thermally Sintered Simulant (TSS)

- Temperatures maintained between -2.5C and -6C
 - Ensure no melting
- 7 day duration
- Reproduce vacuum sintering without a vacuum chamber



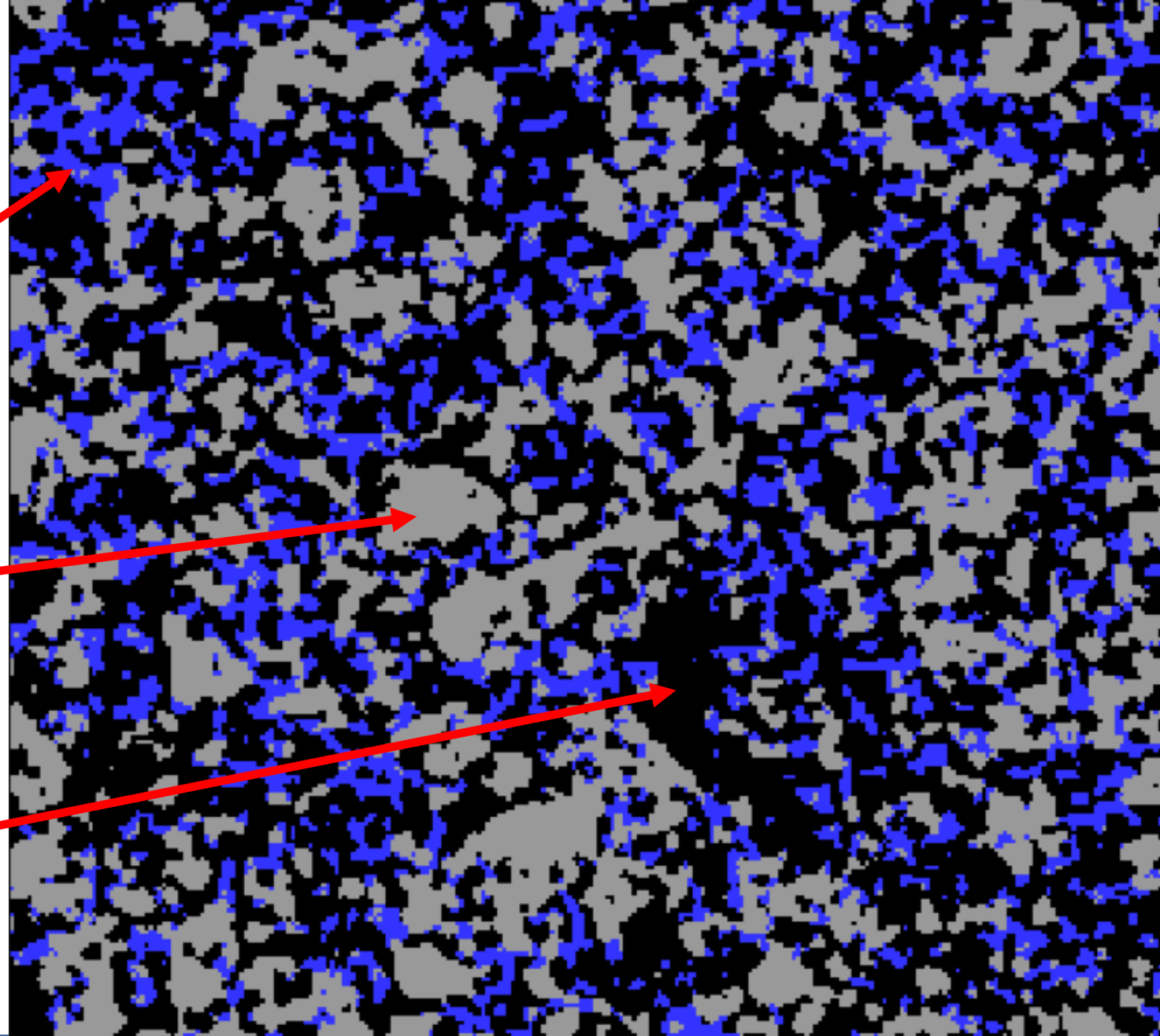
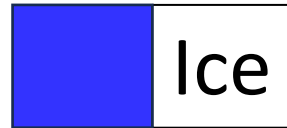
Porosity



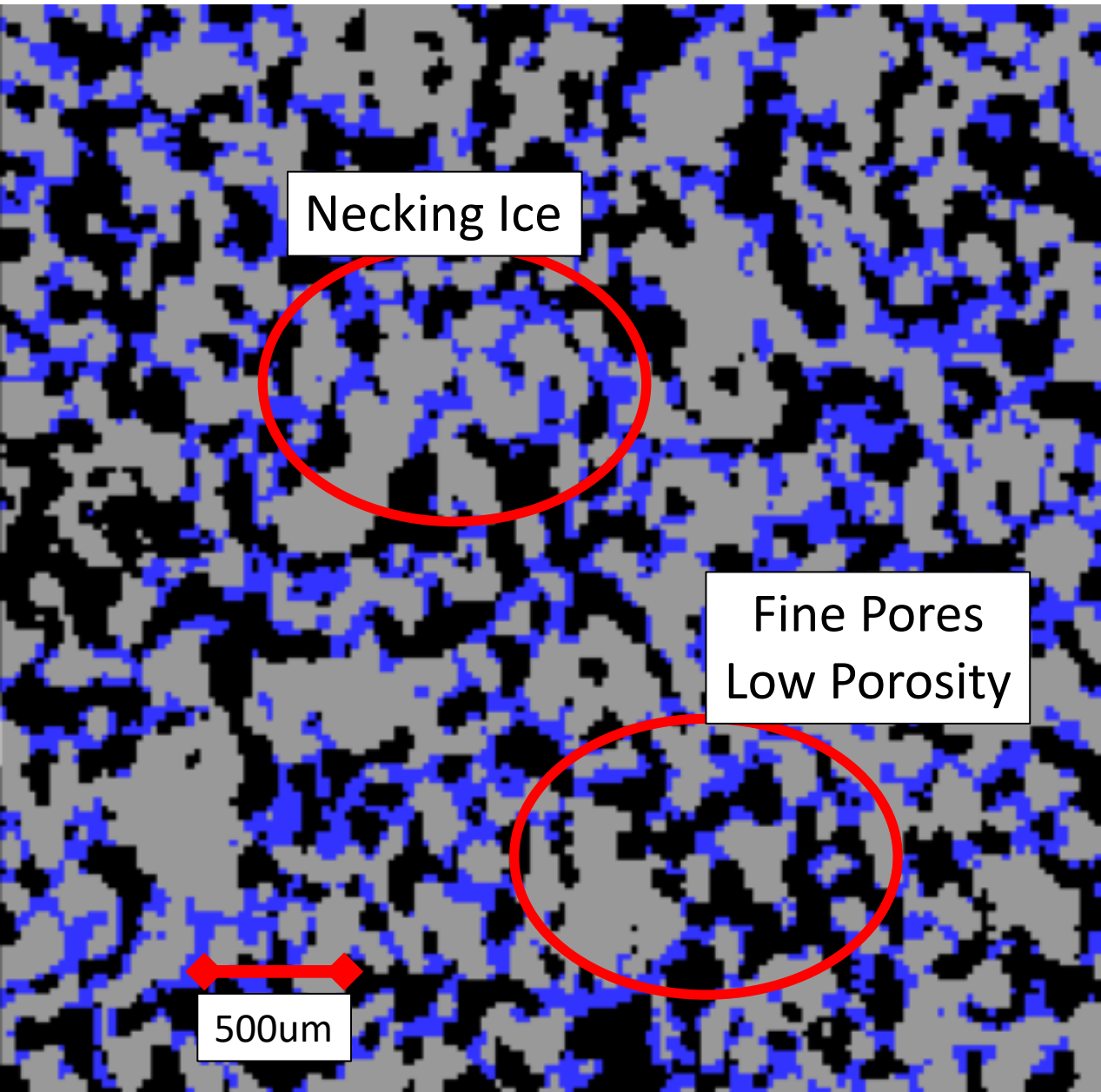
- Dry-mixed ice and regolith cannot densify like wet-mix
 - Even with applied pressure
- Divergent behavior
 - Wet mix decreases in porosity with increased water content

Micro CT Scans Morphology Investigation

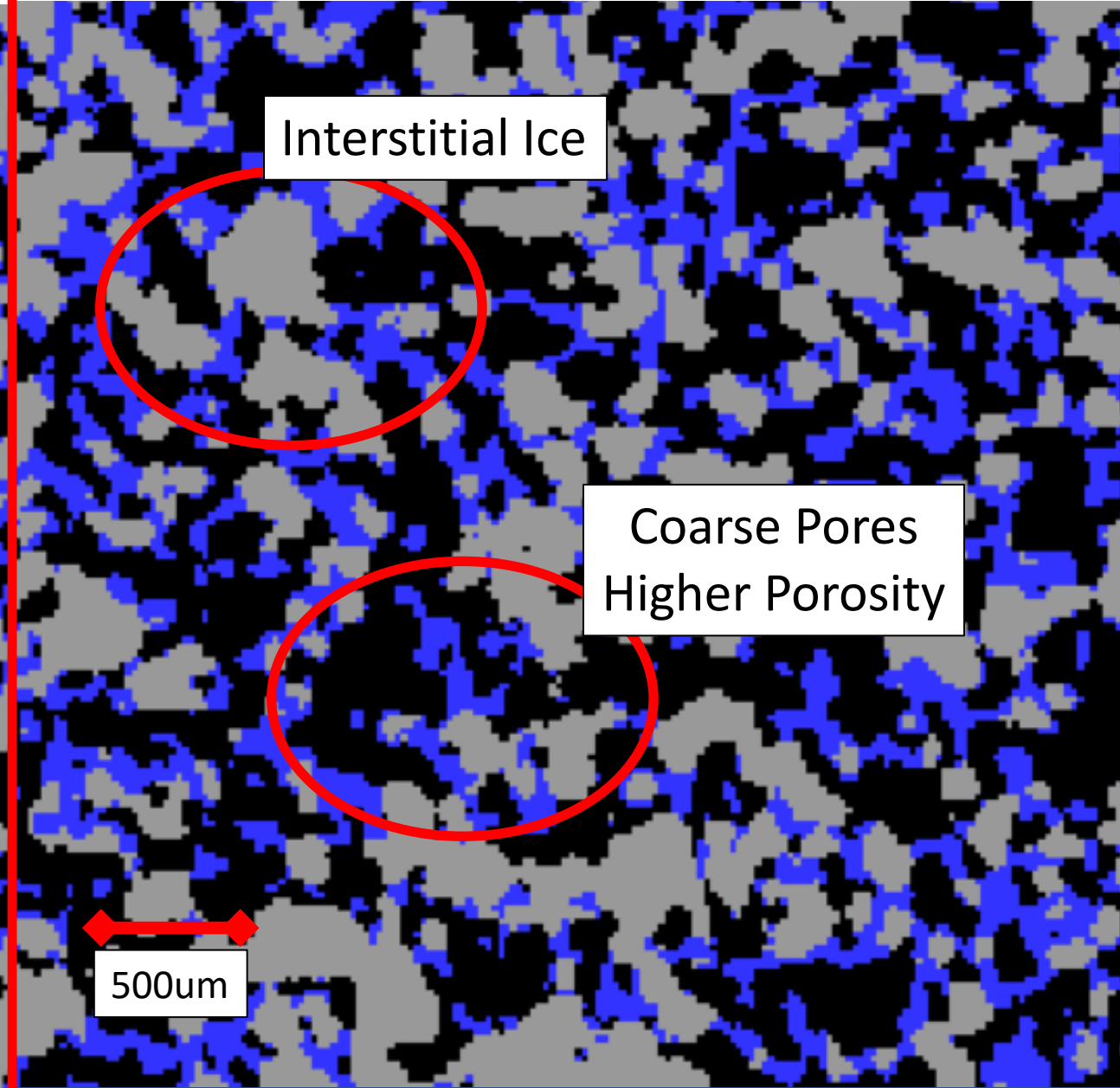
Micro CT Scans



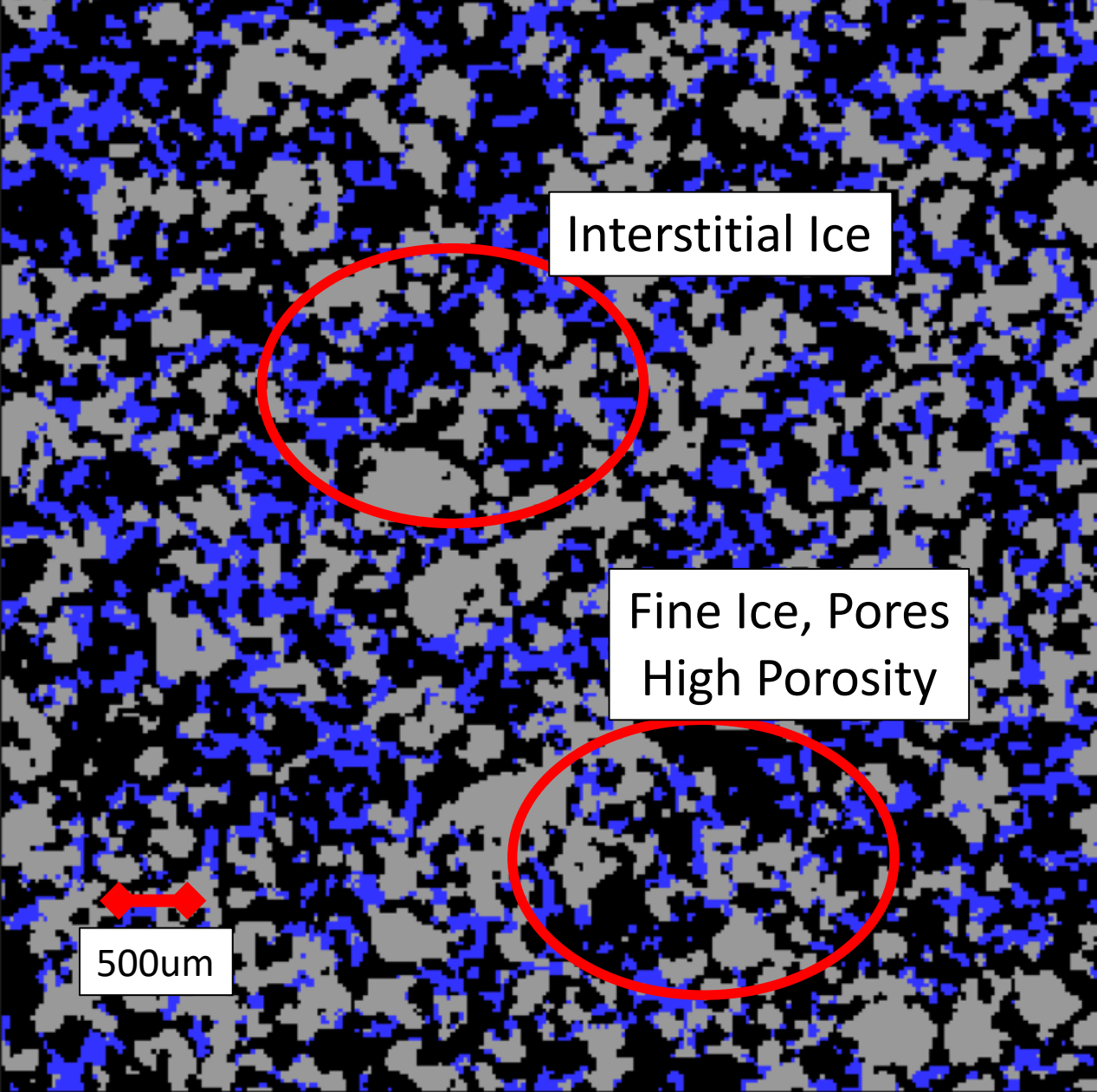
Wet Mix



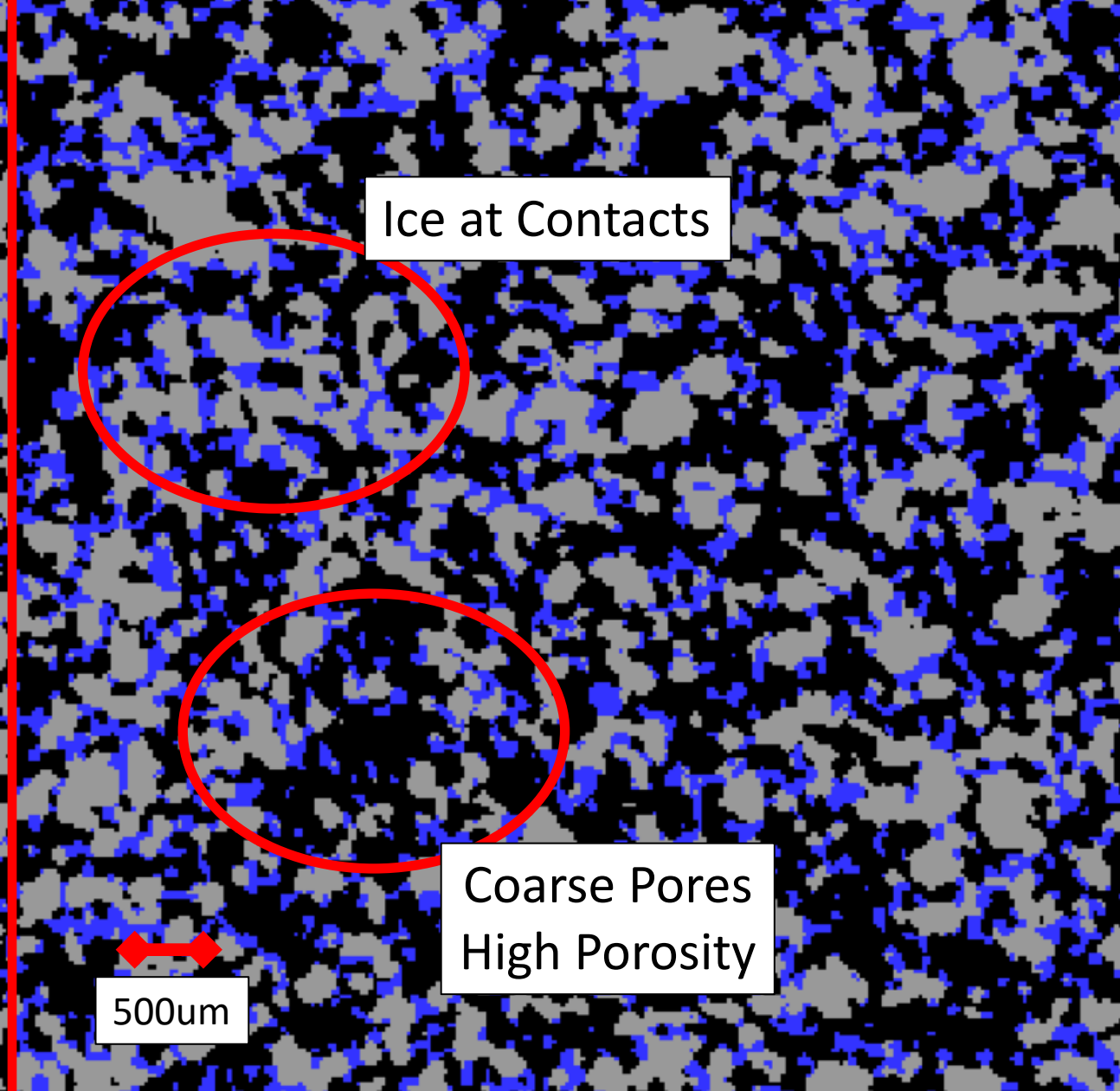
Pressure Sintered



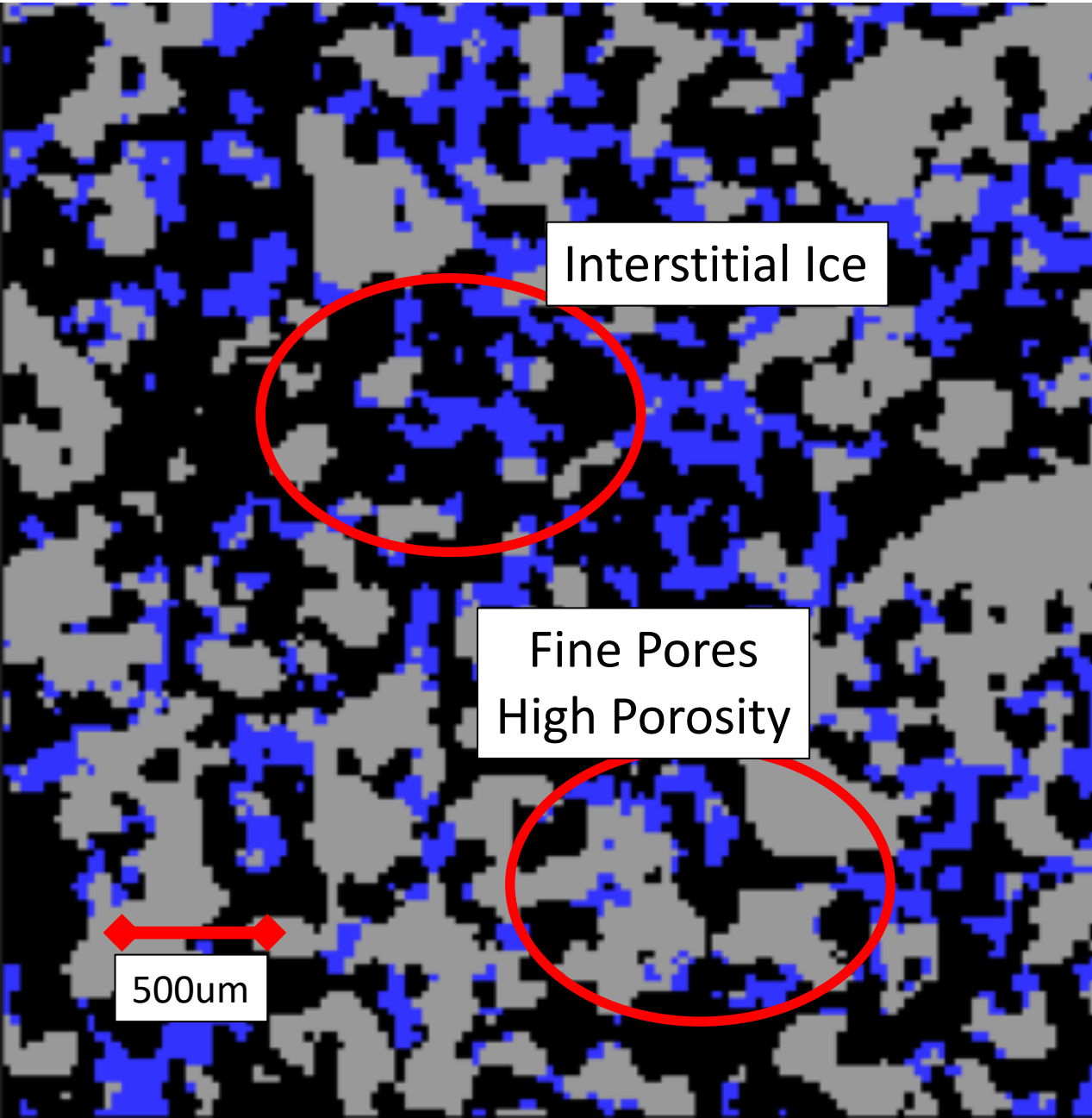
Granular



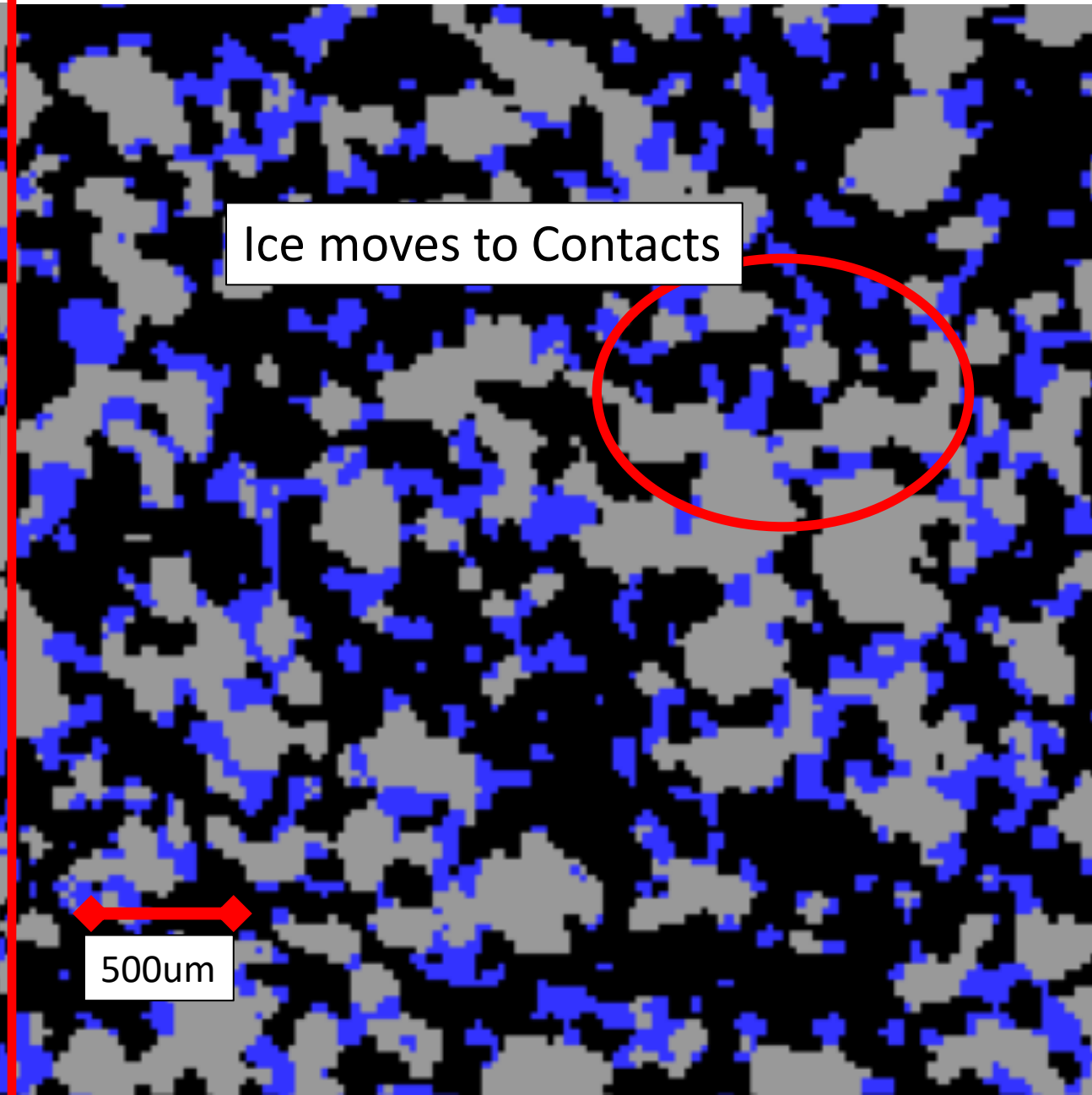
Thermally Sintered



Granular



Thermally Sintered



Conclusions

- Sintering is a rapid, viable volatile modification mechanism on the Moon
- Sintering prevents densification, arrests high porosity (LCROSS)
- New suite of sintered icy regolith simulants available with geologic justification
 - Thermally Sintered, Pressure Sintered, Vacuum Sintered
 - Highly tunable
- Wet mix icy simulant shown to be fundamentally different than sintered
 - Bulk properties: High strength, high density, low porosity
 - Microstructure: Small pores, grain bridging
 - No geologic justification
 - Not conservative, just different

Call to Action

- Prime-1, Viper (Trident, MSolo)
 - Calibrate needed on sintered high-porosity icy regolith simulants
 - Tunable, geologically relevant
- Use sintered icy simulants to develop new excavation, extraction, prospecting, instrumentation technologies

Journal paper,
more to come!



Johnson, D. K., Dreyer, C. B., Cannon, K. M., & Sowers, G. (2024). Pressure Sintered icy lunar regolith Simulant (PSS): A novel icy regolith simulant production method. *Icarus*, 410, 115885.

Works Cited

- Cannon, K. M., & Britt, D. T. (2020). A geologic model for lunar ice deposits at mining scales. *Icarus*, 347. <https://doi.org/10.1016/j.icarus.2020.113778>
- Cooper, P. W. (1997). *Explosives Engineering*.
- Costello, E. S., Ghent, R. R., Hirabayashi, M., & Lucey, P. G. (2020). Impact Gardening as a Constraint on the Age, Source, and Evolution of Ice on Mercury and the Moon. *Journal of Geophysical Research: Planets*, 125(3). <https://doi.org/10.1029/2019JE006172>
- Costello, E. S., Ghent, R. R., & Lucey, P. G. (2018). The mixing of lunar regolith: Vital updates to a canonical model. *Icarus*, 314, 327–344. <https://doi.org/10.1016/j.icarus.2018.05.023>
- Costello, E. S., Ghent, R. R., & Lucey, P. G. (2021). Secondary Impact Burial and Excavation Gardening on the Moon and the Depth to Ice in Permanent Shadow. *Journal of Geophysical Research: Planets*, 126(9). <https://doi.org/10.1029/2021JE006933>
- Davison, L. (2008). *Fundamentals of Shock Wave Propagation in Solids* (L. Davison, Y. Horie, & R. A. Graham, Eds.). Springer.
- Deutsch, A. N., Head, J. W., & Neumann, G. A. (2020). Analyzing the ages of south polar craters on the Moon: Implications for the sources and evolution of surface water ice. *Icarus*, 336. <https://doi.org/10.1016/j.icarus.2019.113455>
- Dreyer, C. B., Abbud-Madrid, A., Atkinson, J., Lampe, A., Markley, T., Williams, H., McDonough, K., Canney, T., & Haines, J. (2018). A new experimental capability for the study of regolith surface physical properties to support science, space exploration, and in situ resource utilization (ISRU). *Review of Scientific Instruments*, 89(6). <https://doi.org/10.1063/1.5023112>
- Gertsch, L., Gustafson, R., & Gertsch, R. (2006). Effect of water ice content on excavatability of lunar regolith. *AIP Conference Proceedings*, 813, 1093–1100. <https://doi.org/10.1063/1.2169290>
- Holsapple, K. A. (1993). THE SCALING OF IMPACT PROCESSES IN PLANETARY SCIENCES. In *Annu. Rev. Earth Planet. Sci* (Vol. 21). www.annualreviews.org
- Gertsch, L. S., Rostami, J. ;, & Gustafson, R. (2008). *Review of Lunar Regolith Properties for Design of Low Power Lunar Excavators Lunar Excavator* (Vol. 2). <https://scholarsmine.mst.edu/icchgehttps://scholarsmine.mst.edu/icchge/6icchge/session10/2>

Works Cited Contd.

- Lucey, P. G., Costello, E., Hurley, D. M., Prem, P., Farrell, W. M., Petro, N., & Cable, M. (2020). *RELATIVE MAGNITUDES OF WATER SOURCES TO THE LUNAR POLES*.
- Lucey, P. G., Petro, N., Hurley, D. M., Farrell, W. M., Prem, P., Costello, E. S., Cable, M. L., Barker, M. K., Benna, M., Dyar, M. D., Fisher, E. A., Green, R. O., Hayne, P. O., Hibbitts, K., Honniball, C., Li, S., Malaret, E., Mandt, K., Mazarico, E., ... Orlando, T. (2022). Volatile interactions with the lunar surface. In *Geochemistry* (Vol. 82, Issue 3). Elsevier GmbH. <https://doi.org/10.1016/j.chemer.2021.125858>
- Melosh, H. J. (1989). *Impact Cratering A Geologic Process* (H. Charnock, J. F. Dewey, S. Conway Morris, A. Navrotsky, E. R. Oxburgh, R. A. Price, & B. J. Skinner, Eds.). Oxford University Press.
- Ong, L., Asphaug, E. I., Korycansky, D., & Coker, R. F. (2010). Volatile retention from cometary impacts on the Moon. *Icarus*, 207(2), 578–589. <https://doi.org/10.1016/j.icarus.2009.12.012>
- Petrenko, V. F., & Whitworth, R. W. (1999). *Physics of Ice*. Oxford University Press.
- Pitcher, C., Komle, N., Leibniz, O., Morales-Calderon, O., Gao, Y., & Richter, L. (2015). Investigation of Properties of Icy Lunar Polar Regolith Simulants. *Advances in Space Research*.
- Prem, P., Artemieva, N. A., Goldstein, D. B., Varghese, P. L., & Trafton, L. M. (2015). Transport of water in a transient impact-generated lunar atmosphere. *Icarus*, 255, 148–158. <https://doi.org/10.1016/j.icarus.2014.10.017>
- Purrington, C., Sowers, G., & Dreyer, C. (2022). Thermal Mining of volatiles in lunar regolith simulant. *Planetary and Space Science*, 222. <https://doi.org/10.1016/j.pss.2022.105550>
- Siegler, M., Aharonson, O., Carey, E., Choukroun, M., Hudson, T., Schorghofer, N., & Xu, S. (2012). Measurements of thermal properties of icy Mars regolith analogs. *Journal of Geophysical Research: Planets*, 117(3). <https://doi.org/10.1029/2011JE003938>
- Șlumba, K., Sargeant, H. M., & Britt, D. T. (2022). Permanently Shadowed Region (PSR) Icy Regolith Simulant. In *Journal of Geophysical Research* (Vol. 813, Issue 8). Planets. <https://exolithsimulants.com>
- Sowers, G. F., & Dreyer, C. B. (2019). Ice mining in lunar permanently shadowed regions. *New Space*, 7(4), 235–244. <https://doi.org/10.1089/space.2019.0002>
- Stewart, B. D., Pierazzo, E., Goldstein, D. B., Varghese, P. L., & Trafton, L. M. (2011). Simulations of a comet impact on the Moon and associated ice deposition in polar cold traps. *Icarus*, 215(1), 1–16. <https://doi.org/10.1016/j.icarus.2011.03.014>
- Stewart, S. T., & Ahrens, T. J. (2005). Shock properties of H₂O ice. *Journal of Geophysical Research: Planets*, 110(3), 1–23. <https://doi.org/10.1029/2004JE002305>

Backup Slides

Impact Gardening with Water Ice

- Secondary impacts 50x more flux than primary impacts
 - Primary ~ 20 km/s \gg 10 GPa
 - Secondary 0.5 - 1.5 km/s: 1-4 GPa
 - Shock testing, ice melts completely at 4 GPa
 - 9 – 50 GPa required for insipient, complete vaporization, respectively
- Cannon et al. (2020)
 - Primary impacts only, no thermal modifications
 - Ice concentration peak ~ 50 cm depth, 10s of cm plentiful
- Costello et al. (2021)
 - Primary + secondary impacts
 - Complete desiccation by secondary impactors
 - Desiccated, 1 Ga = 1m, 3 Ga = 3m

Optimistic

Pessimistic